

RAMPART CRATERS: MECHANISM FOR EXCAVATION OF A MARTIAN BIOTA OF HALOTHERMOPHILIC ARCHAEA-BACTERIA. Steven W. Mitchell, Department of Physics and Geology, California State University, Bakersfield, CA 93311 (smitchell@csbak.edu)

Introduction: Investigations of bacterial phylogeny using nucleic acid sequencing suggest that the first life forms to evolve on Earth were hyperthermophilic archaea and bacteria [1]. This biota probably pursued microbial chemosynthesis based on the anaerobic metabolization of sulfur and/or the aerobic oxidation of hydrogen sulphide, hydrogen, methane or Fe(II) to Fe(III) [2,3]. Such a biota would have evolved and spread rapidly, perhaps within 30 Myr beginning near the end of the early bombardment (approx. 3.8 Bya) [2,4].

A Martian Biota: Conditions similar to those which led to the origin of life on the Earth do seem to have existed on Mars during the late Noachian [5,6]. If life did evolve on Mars, an archaea-bacteria biota may have had the opportunity to diversify and spread, but its biomass would have been relatively much smaller [7]. Potential late Noachian-Hesperian aqueous, endolithic, and subsurface niches would have been present in lacustrine/ocean, fluvial, permafrost, and hydrothermal settings in the context of a probable alkaline hydrosphere [8-12]. Present surface conditions, however, seem to exclude the possibility of even relict exposed or endolithic niches [13,14]. Consequently, subsurface hydrothermal environments have been considered mostly likely to provide niches for a surviving Martian biota [15,16]. By the late Amazonian, a significant decline in volcanism had restricted both the size and extent of potential hydrothermal sources so that the discovery/study of any associated ecosystem may be very difficult [16].

Rampart Craters: Martian rampart craters form when ejecta surges loose fluidizing vapors and the transported material is deposited [17]. Subsurface volatiles in the form of groundice and permafrost extend to a depth of perhaps several kms [18]. If a subsurface viable spore-stage archaea-bacteria biota was present in the groundice, it would be incorporated in the volatilized ejecta (and subsurface plume). A reani-

mation of the spore assemblage would result in a microbial bloom. Following complete cooling of the impact ejecta/subsurface plume a new spore assemblage would be formed. Halophilic bacteria spores can remain viable for at least 250 Myr [19]. Given the apparent Martian cratering rate [20], the long-term survival of a biota based on intermittent hydrothermal and impact-generated events might be possible.

Exobiology Exploration: Rampart craters may provide the most likely locations for the preservation of viable Martian life forms. At shallow depths they may preserve concentrations of halothermophilic archaea-bacteria spores. Rampart craters formed in areas of late Amazonian volcanism would be the most promising locations for the search for spore concentrations.

References [1] Glasby G. P. (1998) *Microbiol. Rev.* 51, 221-271. [2] Woese C. R. (1998) *Episodes*, 21, 252-256. [3] Forterre P. (1995), *Planet. Space Sci.*, 43, 167-177. [4] Oberbeck V. R., et al. (1989) *Origin Life Evol. Biosphere*, 19, 549-560. [5] Oberbeck, V. R. (1990) *LPSC Proc.* 20th, 473-478. [6] Mancinelli R. L. & Banin, A. (1995) *Adv. Space Res.* 15(3), 171-176. [7] Jakosky, B. M. & Shock, E. L. (1998), *JGR* 103, 19,359-19,364, [8] Kempe, S. & Kazmierczak, J. (1997), *Planet Space Sci.* 45, 1493-1499. [9] Thomas, D. J. & Schimel, J. P (1991) *Icarus*, 91 199-206. [10] Carr, M. H. (1996) *Endeavour* 20, 56-60. [11] Scott, D. H., et al. (1991) *Origins Life Evol. Bio.* 21, 189-198. [12] Ostroumov, V. (1995), *Adv. Space Res* 15(3), 229-236. [13] Friedman, E. I. & Koriem, A. M. (1989) 9(6), 167-172. [14] Brack, A. (1996) *Planet, Space Sci.* 44, 1435-1440. [15] Shock, E. L. (1997) *JGR*, 102, 23,687-23,694. [16] Boston, P. J., et al. (1992) *Icarus*, 95 300-308. [17] Wohletz, K. H., et al. (1983) *Icarus* 56, 15-37. [18] Carr, M. H. & Schaber, G. G. (1977) *JGR* 82, 4039-4054. [19] Vreeland, R. H. (1995) *GSA Abs* 27, A305. [20] Haretmann, W. K. (1999) *Meteoritics & Planet Sci* 34, 167-177.